

# Shedding CARMA Light on the Dark Matter in NGC 6503?

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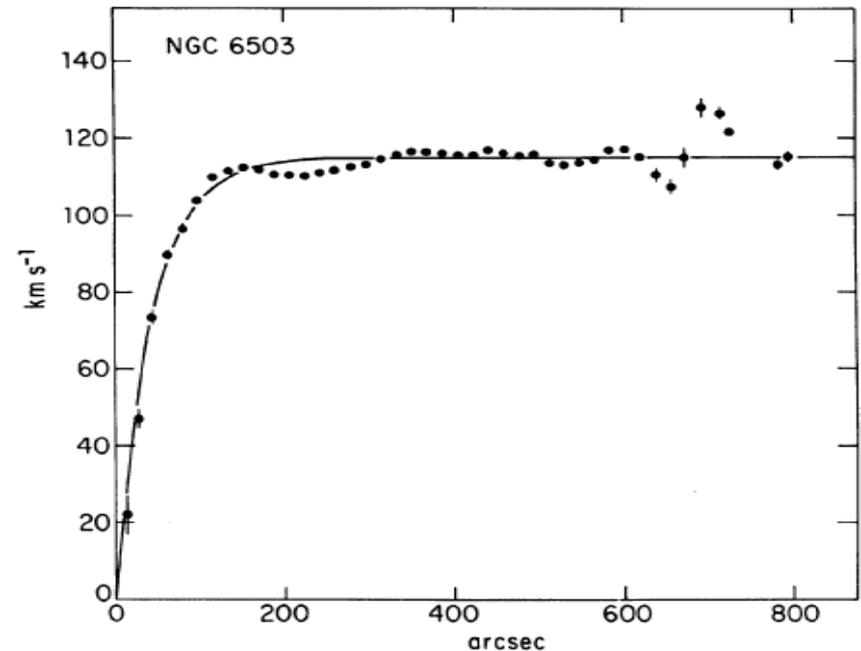
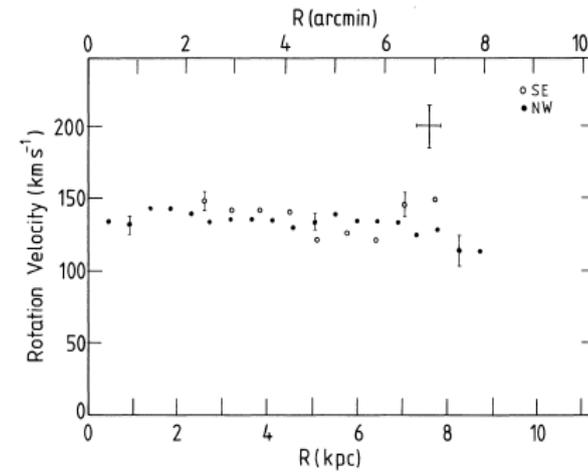
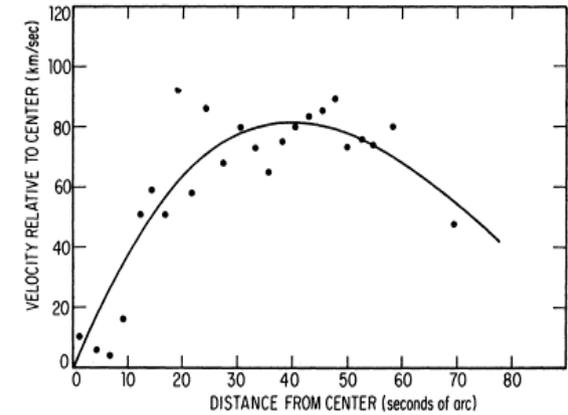
# NGC 6503



- Type: SA(s)cd
- RA/DEC:  $17^{\text{h}}49^{\text{m}}$   $+70^{\text{d}}09^{\text{m}}$
- VSYS: 28 km/s (HEL) 40 km/s (LSR)
- Distance: 5.2 Mpc or 6 Mpc (old: 3.7 Mpc)
- Scale:  $1'' \approx 18$  pc,  $1' \approx 1$  kpc
- Size: 5.4 kpc (Optical) 9.0 kpc (HI)
- Shape:  $74^{\circ}$  (incl)  $121^{\circ}$  (position angle)
- Luminosity: -17.68 or  $1.5 \cdot 10^9 L_{\odot}$  (in B band)
- Mass:  $1.6 \cdot 10^9 M_{\odot}$  (in HI)

# Old Data Sets

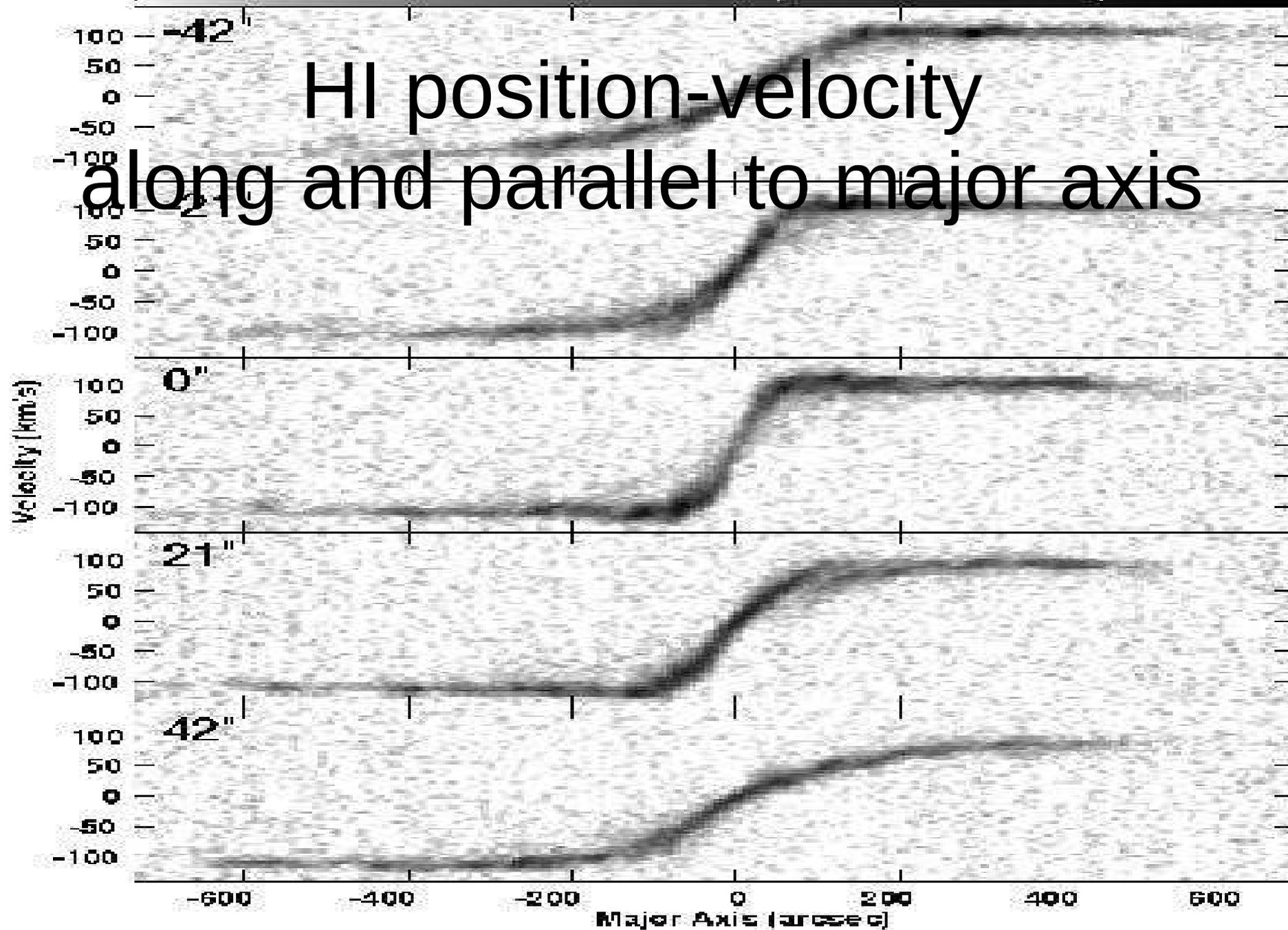
- Long slit spectra (Burbidge et al. 1964), out to 60"
- HI synthesis WSRT (Shostak et al. 1981), out to 500", **50" resolution**
- HI synthesis VLA (v.Moorsel & Wells 1985), out to 800" **15x31" resolution**



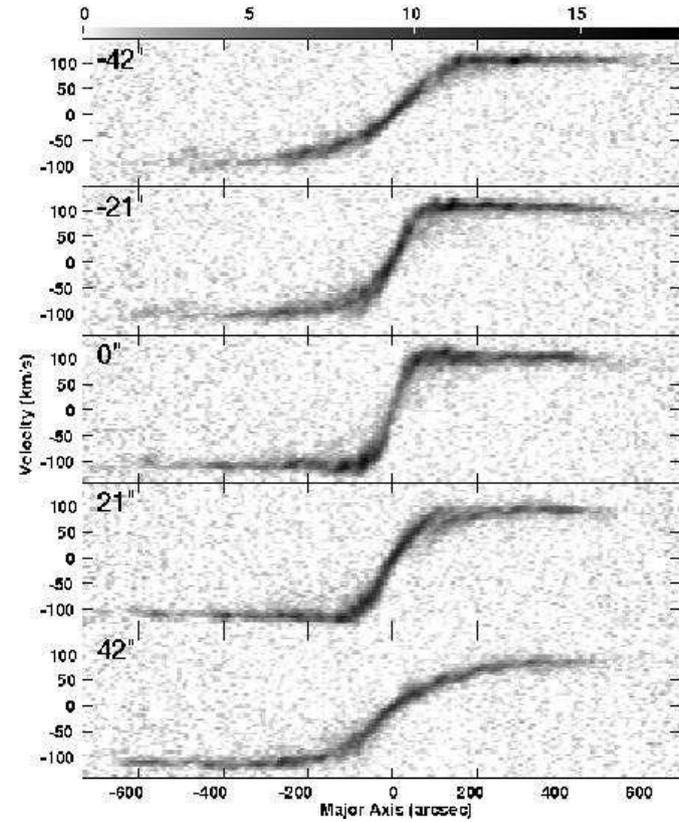
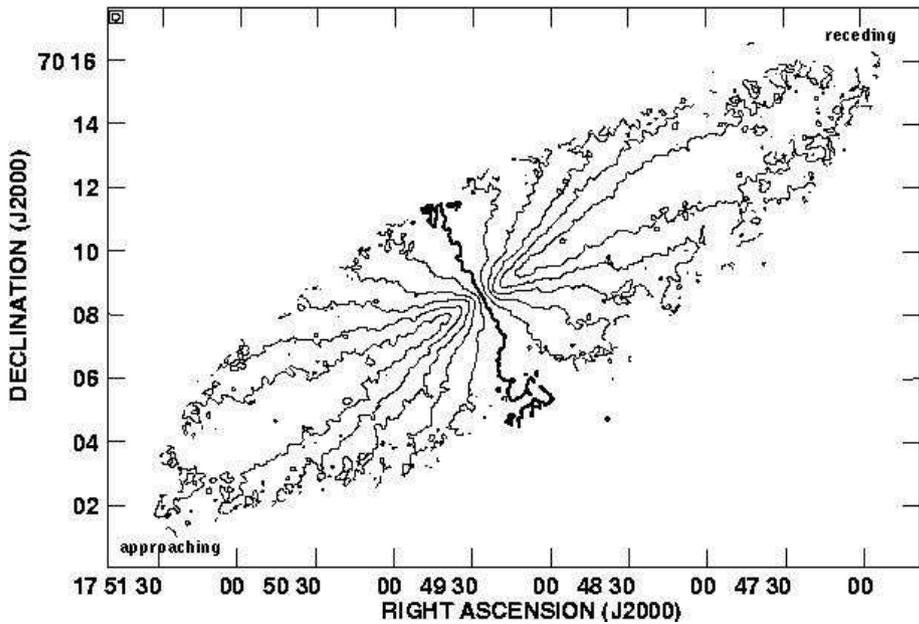
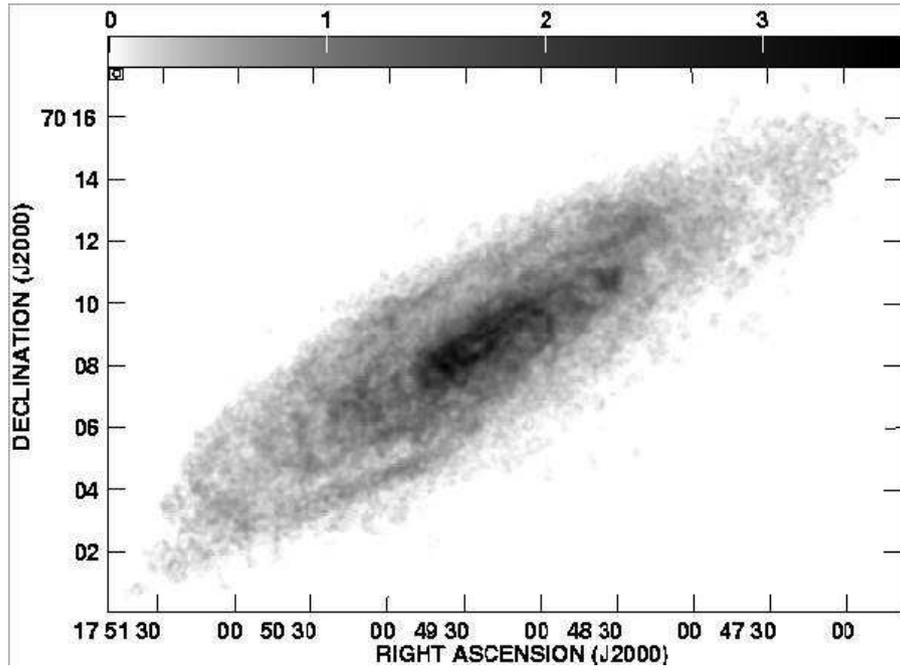
# New Data Sets

- HI synthesis VLA (1996/2009), out to 900", **14" resolution**
- H $\alpha$ , DensePak, central 120", **4"**
- CO (1-0), central 100" CARMA synthesis, **5"**

# HI position-velocity along and parallel to major axis

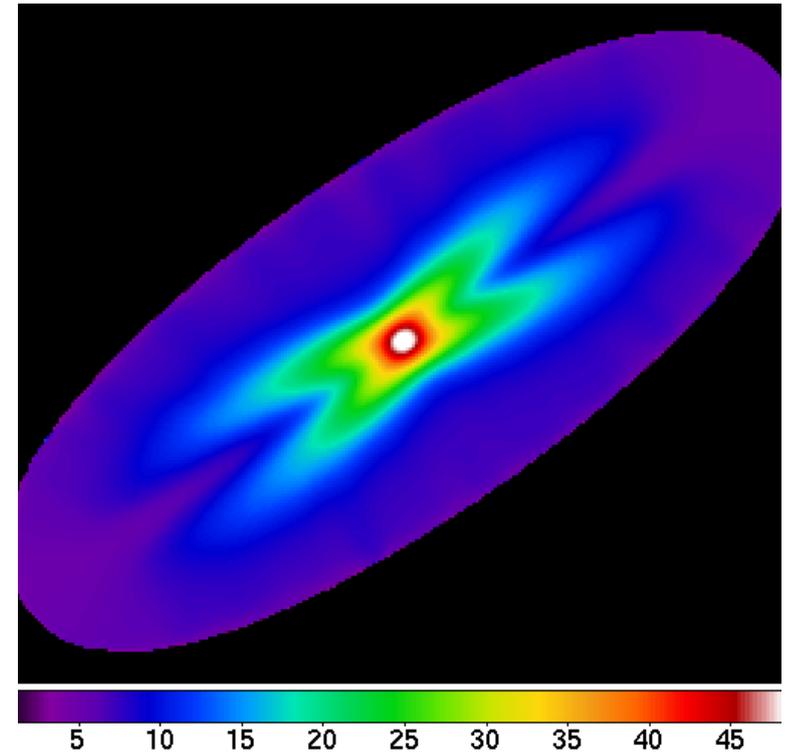
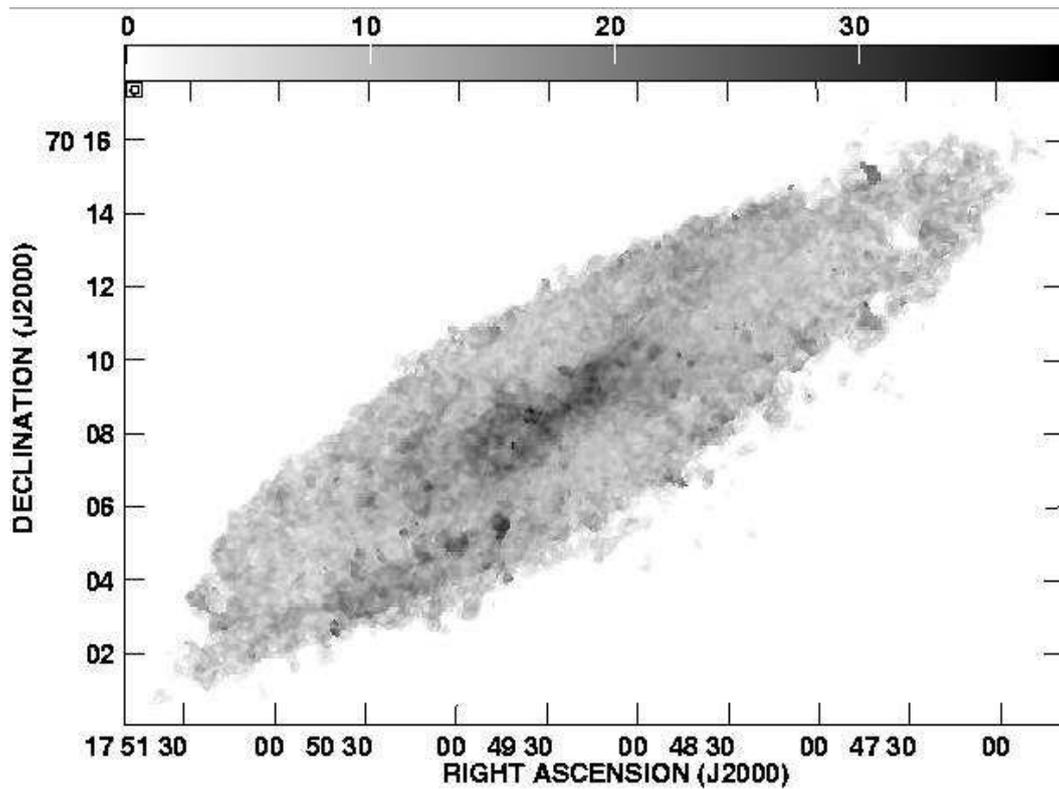


# VLA HI at 14"

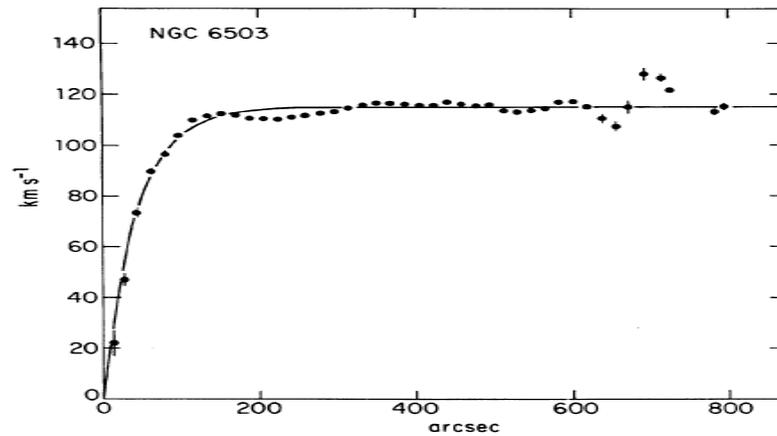
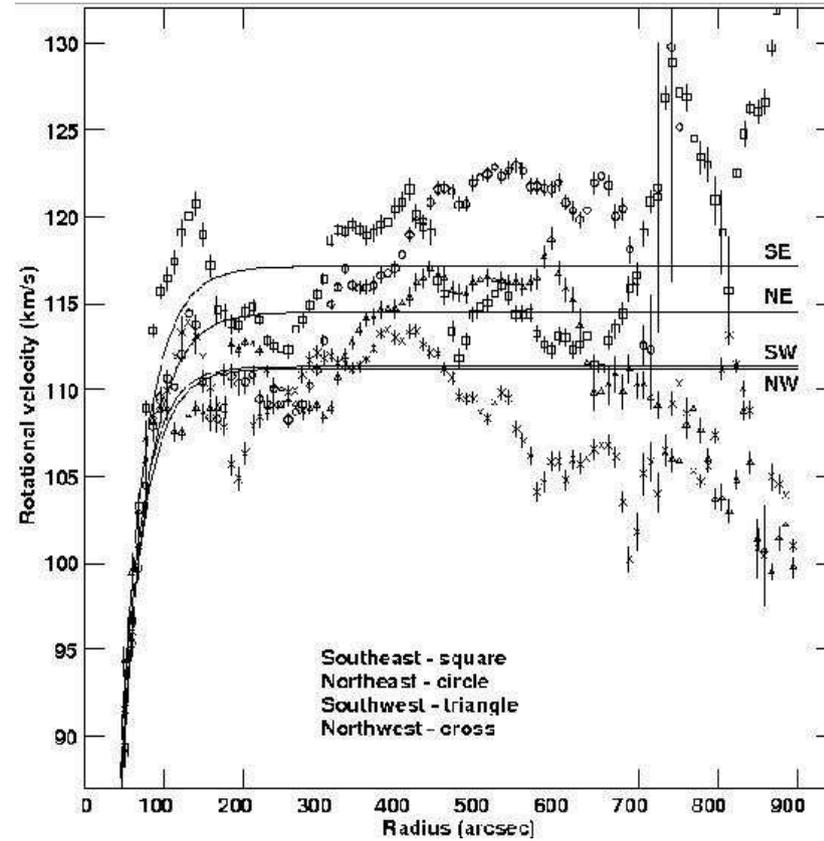


# HI – velocity dispersion ( $\sigma$ )

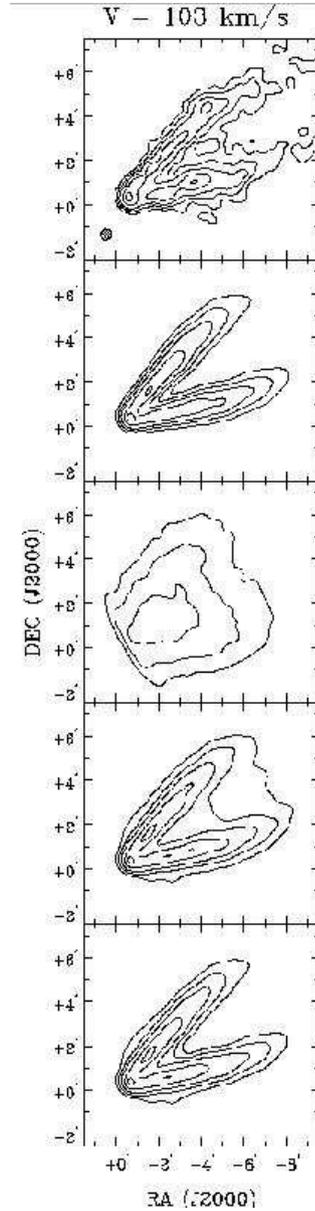
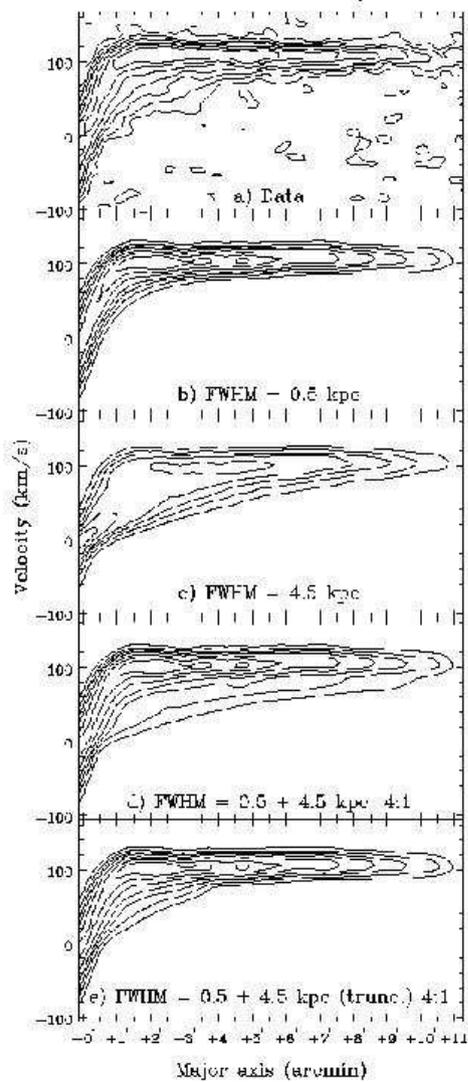
- Observations
- Model w/ constant  $\sigma$



# “Rotation Curve”



# Modeling

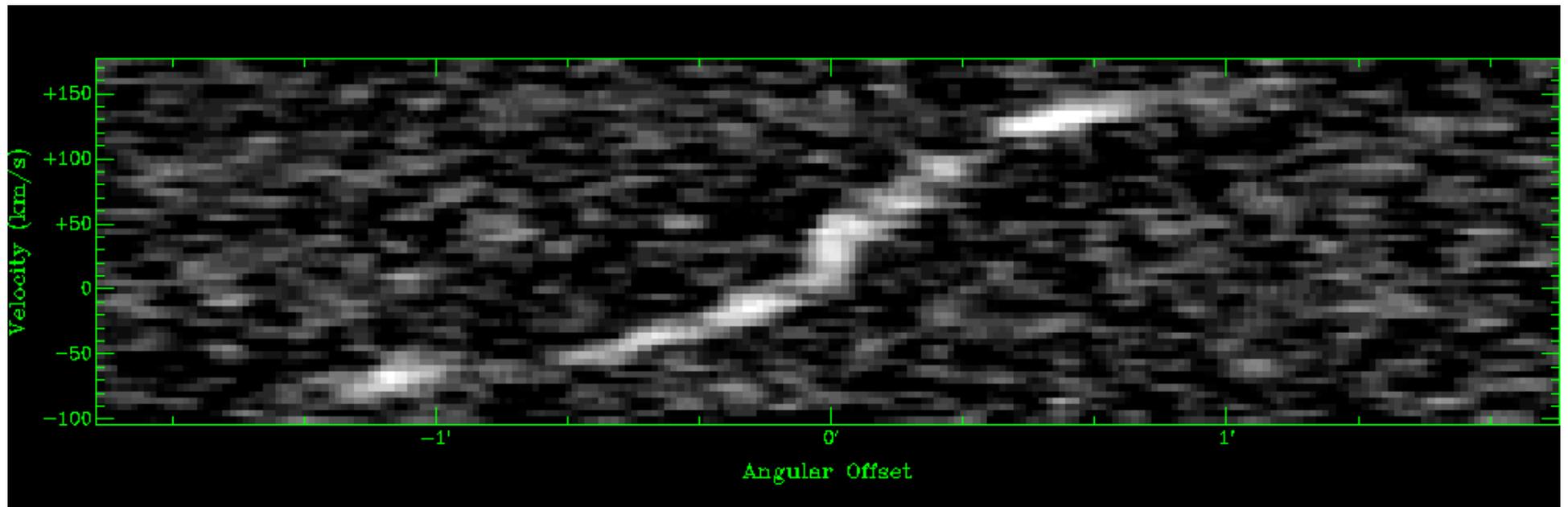
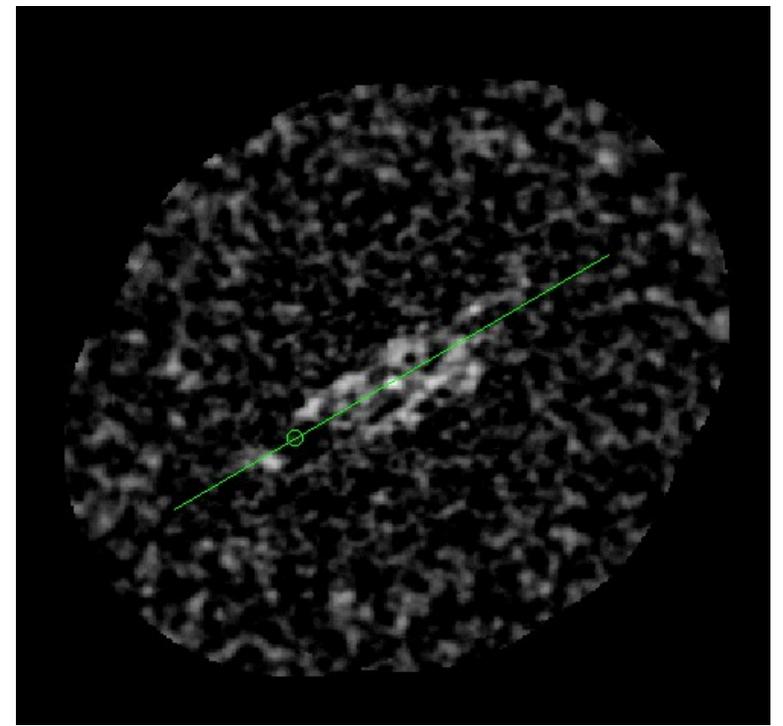


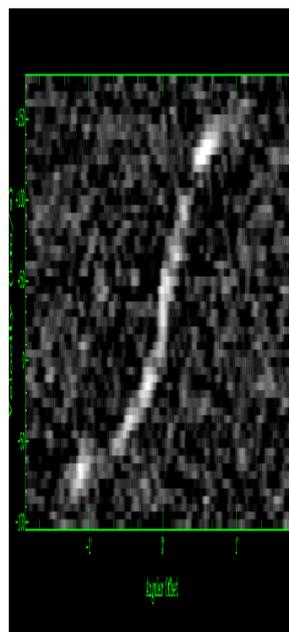
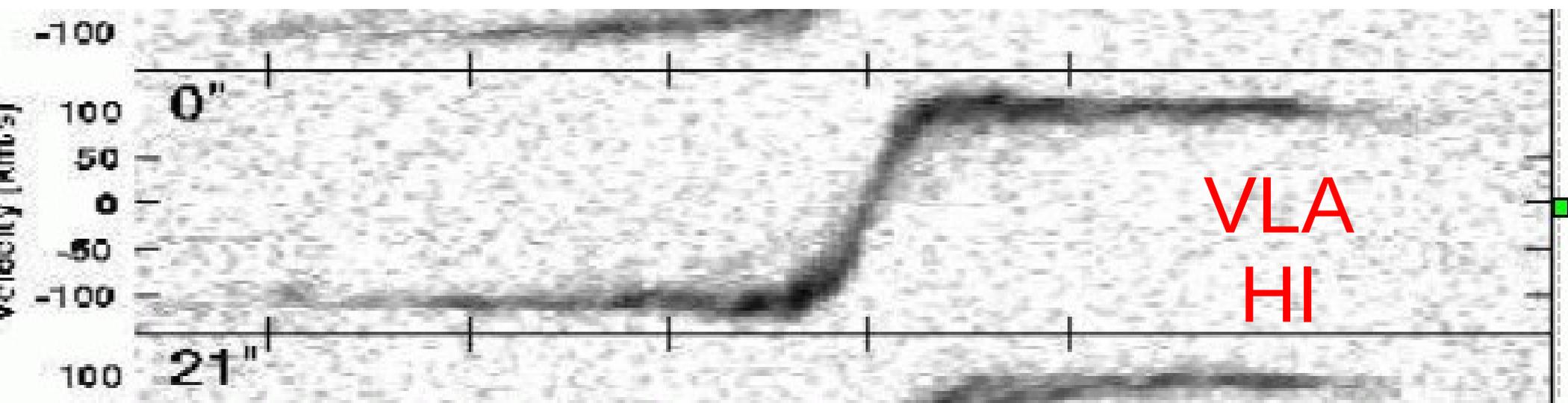
# CARMA

- 23 antenna: 6 OVRO + 9 BIMA + 8 SZA (new!)
- Synthesis array, up to 2 km baselines
  - Resolution 0.3'' ... 10'' (at 3 mm)
- Frequency bands:
  - 26-36 GHz (1 cm)
  - 80-115 GHz (3mm) e.g. CO (1-0)
  - 215-270 GHz (1mm) e.g. CO (2-1)



# CARMA CO(1-0): Position-Velocity along major axis

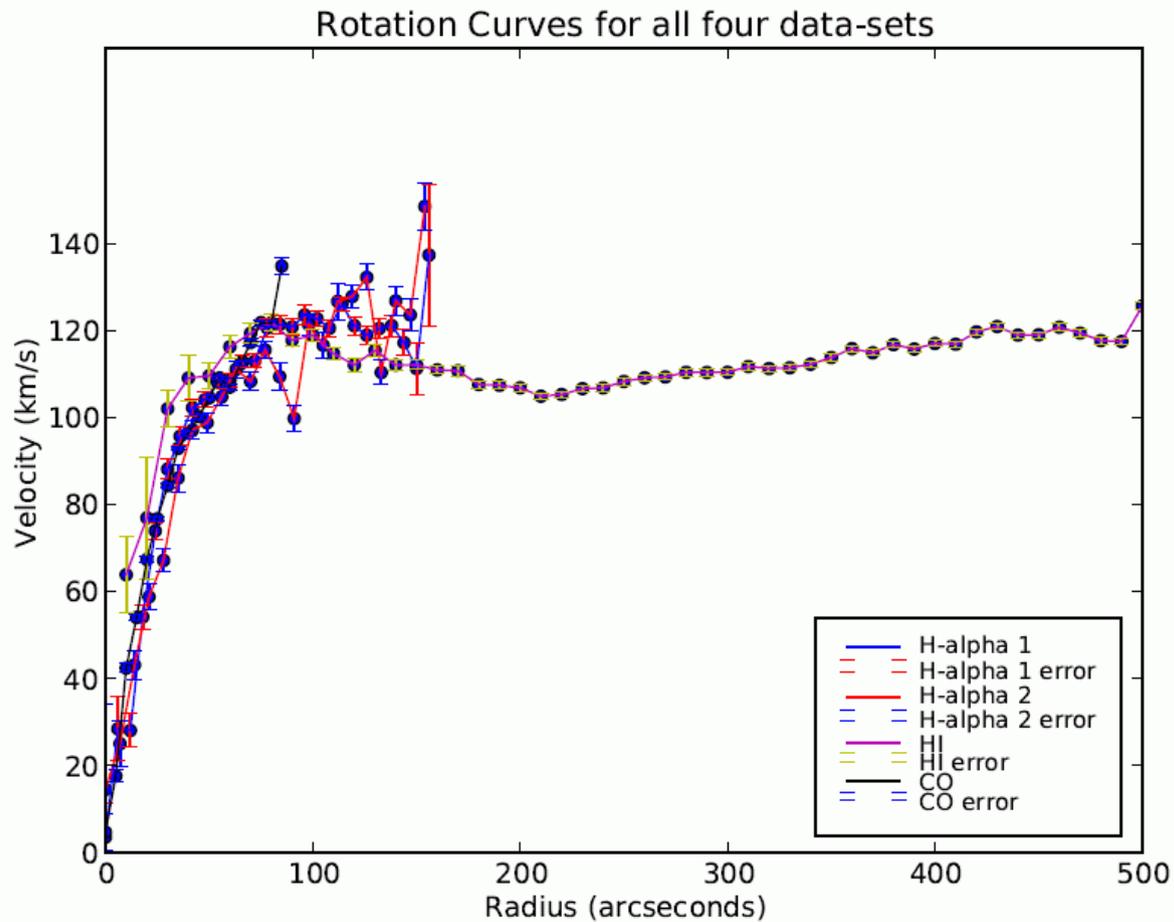




# Data Cube to Rotation Curve

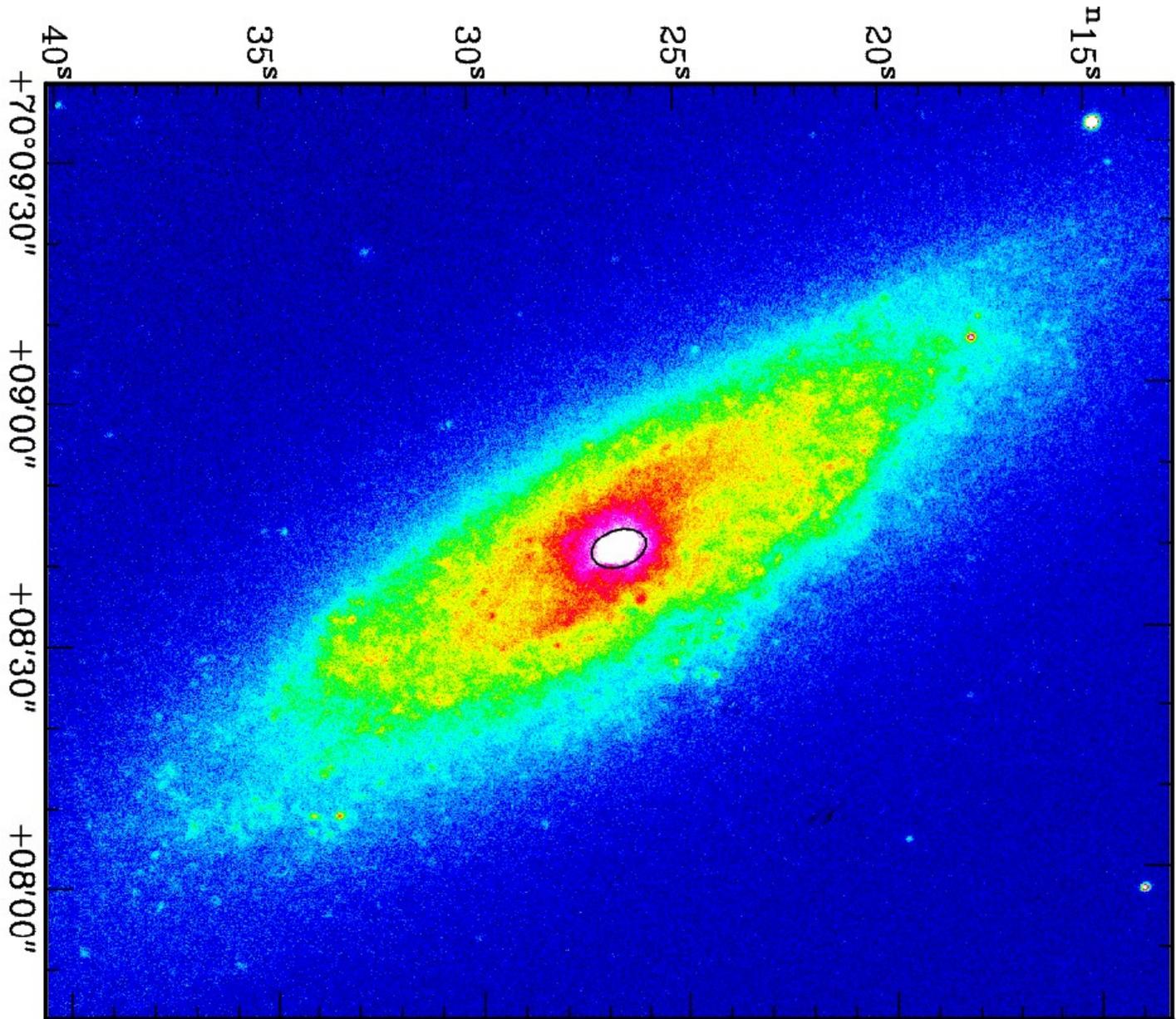
- Moment analysis
  - Classically:  $I(x,y,vz) \rightarrow V_z(x,y) \rightarrow V_{\text{rot}}(r)$
  - Lots of issues at each step:
    - Moments vs. profile fitting vs. edge-trace
    - Beam smearing: spatial and velocity
    - Optically thin or thick tracer + coverage
- Full 3D modeling
  - Fairly new, and expensive; iterative (e.g. Swaters)

# Rotation Curve comparison



# H-band: a nuclear bar?

(Freeland et al. 2009 submitted)



# Unresolved (5") high velocity gradient in the central 100 pc

- Bar at favorable viewing angle?
  - Strong bars don't form in hot stellar systems?
  - Unusually small bar (but bars can be nested)
- DM Cusp?
- DM around SMBH?
  - But very small bulge, so very small central BH?
- Star Formation?
- Something else?

# Future Work

- High res (1.5" or 0.3") of nuclear region
  - CARMA C-array starts April 6<sup>th</sup> !
- Full modeling, perhaps with a nuclear bar